# Vacuum Ultraviolet (VUV) Photoabsorption Spectroscopy with Laser Induced Plasmas

EMSLIBS

#### John T. Costello School of Physical Sciences & NCPST, DCU



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# Talk Outline



Ollscoil Chathair Bhaile Átha Cliath Dublin City University

- DCU Laser Plasma & AMO Group
- Dual Laser Plasma Photoabsorption (DLPP) The Idea
- Why Do (Short Wavelength) Photoabsorption?
- Laser Plasma Continua History & Physical Origin
- DLPP Some History
- Some Recent VUV/EUV DLPP Results on Atoms and Molecules
- "For What's Next" Perspectives for X-VUV....
- Acknowledgements



# LP & AMO Physics Group at DCU (Pls)



Ollscoil Chathair Bhaile Átha Cliath Dublin City University John Costello (LP Spectroscopy & Free Electron Lasers) Emr. Eugene Kennedy (LP Spectroscopy & Synchrotrons) Jean-Paul Mosnier (Synchrotrons) Lampros Nikolopoulos (AMO Theory) Paul van Kampen (Synchrotrons/ Physics Education) Current Postdoctoral and PhD students:

Lazaros Varvarezos, James Campbell, Séamus Cummir

Lazaros Varvarezos, James Campbell, Séamus Cummins, Stephen Durkan, Andrew Foremski, Ross McGarry, Adam Prior and Sadaf Syedah Zehra.

**Recent PhD Graduates:** 



2021: Muhammad Bilal Alli, Hu Lu & Tejaswi Katravulapally 2020: Ben Delaney, Stephen Davitt & Lazaros Varvarezos,

# LP & AMO Physics Group at DCU (PIs)



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#### **Current Projects:**

- 1. VUV & EUV Photoabsorption via the Dual Laser Plasma Technique
- 2. UV/VUV LIBS (Steel, Pharma & Wind Turbine Blades (SC))
- 3. Photoionization of Atoms and Molecules in EUV FEL Fields
- 4. Ultrafast FEL Pump EUV Probes of Atoms and Molecules
- 5. Photoionization of Small Molecular lons at Synchrotrons (ETK)
- 6. Development of an Ultrafast Laser Spectroscopy Laboratory

#### References

- 1. Near-Threshold Two-Photon Double Ionization of Kr in the Vacuum Ultraviolet, L Varvarezos et al , Physical Review A 103 Art. No. 022832 (2021)
- Time-integrated and time-resolved VUV LIBS: a comparative study, S S Zehra, J T Costello, P Nicolosi and P Hayden, Proceedings of SPIE, 10674, 106741H (2018)
- 3. Clocking Auger electrons, D C Haynes et al, Nature Physics 17 pp.512–518 (2021)
- 4. The 5d → 6p EUV photoabsorption spectra of Pb II and Bi III: evidence of excited states, H Lu, L Varvarezos, M Bilal Alli, P Nicolosi, J T Costello and P Hayden, J. Phys. B: At. Mol. Opt. Phys. 53 115001 (2020)
- 5. Evolution of L-shell photoabsorption of the molecular-ion series SiH<sup>n+</sup> (n=1,2,3): Experimental and theoretical studies, E T Kennedy *et al* Phys, Rev. A 97 043410 (2018)
- 6. https://www.dcu.ie/research/ultrafast-spectroscopy / http://uf-dynamics.org

## Dual Laser Plasma Photoabsorption (DLPP) - The Idea

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- Use two (normally) electronically synchronised LPPs
- One acts as the sample of atoms, ions or molecules
- The other acts as a source of continuum radiation
- The continuum emission is usually in the Vacuum-UV (VUV) or the Extreme UV (EUV) spectral region





## Vacuum-UV and Extreme-UV Spectral Ranges





 $E(eV) = (hc/\lambda) = (1239.8)/\lambda(nm)$ 







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# Why Do Photoabsorption?

- Access to ground state (non-emitting) atomic and molecular species in the sample (vapour, plasma, etc.)
- You can then also detect (and potentially quantify) metastable state species
- Photon excitation => electric dipole excitation => less cluttered and quite tractable (from a theoretical perspective) spectra





# Why Do X-VUV Photoabsorption?

## Why specifically at X-VUV photon energies?

- Access to more highly charged ions
- Photoionization continua
- Inner-shell/multi-electron excitations

## Data relevant to:

- Astrophysical spectra and models
- Laboratory plasma modelling & diagnostics
- Fundamental many-body theory
- Plasma/atomic X-ray laser schemes
- MCF & ICF
- DLP data guides large scale synchrotron expts



Ollscoil Chathair Bhaile Átha Cliath Dublin City University Quick reminder - Radiative Processes in Laser Plasmas

1. Bound - Bound Transitions - Line Absorption/Emission

$$A^{n+} + h_V \leftrightarrow A^{n+*} \text{ or } A^{n+*} \rightarrow A^{n+} + h_V'$$

2. Bound - Free Transitions - Recombination/Photoionization\*

 $A^{(n+1)+} + e \leftrightarrow A^{n+} + hv$ 

3. Free - Free Transitions - Bremsstrahlung/Inv. Bremsstrahlung

$$A^{n+} + e(T_1) \leftrightarrow A^{n+} + e(T_2) + h\nu, T_1 > T_2$$





*Free-Free and Free-Bound* processes yield continuum emission spectra suitable for application in absorption spectroscopy





Short wavelength continua emitted from laser produced rare-earth (and neighbouring element) plasmas are predominantly free-bound in origin











Fig. 1. (a) Absorption spectrum of xenon from 80 to 200 Å. The xenon pressure in the spectrograph was 0.05 Torr, and the number of laser pulses used was 30. For details of the xenon spectrum in this region see Madden and Codling.<sup>5</sup> The unmarked weak lines near 200 Å are due to 0 v. Oxygen present in the target gives rise to some emission lines as well. (b) The ytterbium continuum from 60 to 100 Å. The number of laser shots was 20. As in (a), the spectrum was obtained on a Kodak SC5 plate.

P K Carroll et al., Optics Letters 2, pp72-74 (1978)

E T Kennedy et al., Optical Engineering 33, pp3894-3992 (1994)



**Fig. 2** (a) EUV emission spectrum of an aluminum oxide plasma showing the predominance of lines from  $O^{4+}$  in the  $54 \rightarrow 64$  eV photon energy range. (b) and (c) Continuum emission from a tungsten plasma in the 30- and 140-eV spectral ranges.





Motivation: Measure ions of astrophysical interest, tests of databases (e.g., Opacity, etc.)

BR TM Slit RCC

P Recanatini, P Nicolosi & P Villoresi, Phys. Rev. A **64**, Art. No. 012509 (2001)



Spaced resolved emission from a W plasma in the VUV around (a) 49 nm and (b) 69 nm



Normal Incidence DLP Setup



## But why is no line emission observed ?

Line emission is due to complex 4d-4f transition arrays in (typically) 7 - 20 times ionized atoms:

```
4d^{n}5s^{q}5p^{s}4f^{m} \rightarrow 4d^{n-1}5s^{r}5p^{t} 4f^{m+1}, q+s = r+t
```

Furthermore 4f/5p and 4f/5s degeneracy and level crossing gives rise to overlapping bands of low-lying configurations, most of which are populated in the ca.10 - 50 eV plasma

Result - the summed oscillator strength for each 4d - 4f (XUV) and 5p - 5d (VUV) array is spread out over a *supercomplex* of transitions producing bands of unresolved pseudo continua (so called Unresolved Transition Arrays - 'UTA') superimposed on the background continuum



In addition, strong emission lines from simple 4f - 4f transition arrays, e.g., 4d<sup>10</sup>-4p<sup>9</sup>4f in Xe-like ions, are washed out by plasma opacity





There are up to 0.5 million allowed transitions (in LS coupling) over the ~10 eV bandwidth of a UTA

In fact, this is a lower bound since many additional LS forbidden transitions are 'switched on' by the breakdown in LS coupling here - G O'Sullivan et al., J.Phys.B **32**, 1893 (1999)



## Brief Early History/ Highlights of Laser Plasma High-Z/Rare-Earth Continua

- 1. First report of line free continua P K Carroll et al., Opt. Letts. 2, 72 (1978)
- 2. First full study/ applications P K Carroll et al., Appl. Opt. 19, 1454 (1980)
- 3. VUV Radiometric Transfer Standard G O'Sullivan et al., Opt. Letts. 7, 31 (1982)
- 4. Absolute Calibration with Synchrotron J Fischer et al, Appl. Opt. 23, 4252 (1984)
- 5. Photoelectron Spectroscopy Ch. Heckenkamp et al., J. Phys. D 14, L203 (1981)
- 6. First Study for EUV lithography D J Nagel et al., Appl. Opt. **19**, 1454 (1980)
- 7. EUV Reflectometer S Nakayama et al., Physica Scripta **41**, 754 (1990)
- First Industrial Application DuPont Insulator Band Structure
   VUV Reflectance Spectroscopy R H French, Physica Scripta 41, 404 (1990) System subsequently made available commercially from ARC

For a review of the early years including applications in photoabsorption spectroscopy see :

- 1. J T Costello et al., Physica Scripta **T34**, 77 (1991)
- 2. P Nicolosi et al., J. Phys. IV 1, 89 (1991)



- Ease of Production
- Ease of Location
  - Purity (Spectral)
- Wide Spectral Coverage  $(4 \rightarrow 200 \text{ nm})$
- Small Emitting Size (almost point-like, radiography & microscopy)
- Short Pulse Duration (< 100 ps 50 ns)\*
- Easy Synchronisation (Optical or Electro-Optic)
- Insensitivity to ambient pressure
- Shot to shot intensity reproducibility  $\leq 5\%$
- ~10<sup>14</sup> Photons/pulse/sr/nm\*\*

\*Depending on exciting laser pulse

\*\*J Fischer et al, Appl. Opt. 23, 4252 (1984) – 800 mj Nd:YAG plasma generation

### Economy, Ease of Use & Versatility



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FIG. 2. Transmittance T and absorption  $\mu x$  of extreme uv (98 Å) through the plasma. d is the distance from the observed zone to the hottest part, situated at about 0.5 mm from the target.  $T_1$ , transmittance measured with a 12-nsec time lag;  $T_2$ , transmittance measured with a 27-nsec time lag;  $I_E$ , intensity of the continous spectrum emitted by the plasma at the same wavelength.

Doubly Excited Autoionization Resonances in the Absorption Spectrum of Li<sup>+</sup> Formed

in a Laser-Produced Plasma\*



EXTREME-ULTRAVIOLET CONTINUUM ABSORPTION

BY A LASER-GENERATED ALUMINUM PLASMA

P. K. Carroll and E. T. Kennedy Physics Department, University College Dublin, Dublin, Ireland (Received 3 February 1977)



FIG. 1. Schematic diagram of experimental arrangement.





FIG. 2. One-electron of  $Li^+$ . The scale in the vertical (x) direction measures the distance from the target surface.



76.98 82.49 49 82.49

FIG. 3. Two-electron absorption of Li<sup>+</sup>.

1. J T Costello et al., Physica Scripta **T34**, 77 (1991) 2. P Nicolosi et al., J. Phys. IV **1**, 89 (1991)



MSL

#### VUV Photoabsorption of carbon ions - from P Nicolosi et al. (Padua)

Motivation: Measure ions of astrophysical interest, tests of databases (e.g., Opacity, etc.)



#### Normal Incidence DLP Setup

P Recanatini, P Nicolosi & P Villoresi, Phys. Rev. A **64**, Art. No. 012509 (2001)



C<sup>+</sup>. 1.2 J on target in line focus: 9 mm X 0.01 mm / Delay = 58 ns



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J Hirsch, E T Kennedy, J T Costello, L Poletto & P Nicolosi Rev. Sci. Inst. 74, 2992 (2003)



#### **Motivation**



1.Direct imaging of light emitted by a plasma using gated array detectors (e.g., I-CCD) provides information on excited species only

2.Probing plasma plumes using tuneable lasers provides information on non-emitting species but is limited to wavelengths > 200 nm or so

Idea: Pass a *collimated VUV beam* through the plasma sample and measure the spatial distribution of the absorption.



Ollscoil Chathair Bhaile Átha Cliath Dublin City University Why a pulsed, tuneable and collimated beam ?

#### Pulsed

Automatic time resolution: the VUV pulse duration ~ laser pulse duration (~1 - 100 ns)

#### • Tuneable

Can access resonance lines of all atoms & moderately charged ions with resonances between 30 nm and >120 nm Can tune into photoionization continua (cross sections known)

#### Collimated



Can place the sample and CCD anywhere along the beam

#### **VUV Photoabsorption Imaging**

**Experimental Schematic** 



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#### **VUV Photoabsorption Imaging**







Time resolution: ~10 - 30 ns 0 - 10 µsec Inter-plasma delay range: Delay time jitter:  $\pm 1$ ns Acton<sup>™</sup> VM510 (f/12, f=1.0 m) Monochromator: 10 - 35 eV (35 – 120 nm) VUV range: VUV bandwidth: 0.025 eV @25 eV (50mm/50mm slits) ~0.05 nm @ 50 nm **Detector**: Andor<sup>™</sup> BN-CCD, 2048 x 512/13 µm x 13 µm pixels ~120 µm (H) x 150 µm (V) Spatial resolution:

**VUV** Photoabsortion Imaging - Specifications





Ollscoil Chathair Bhaile Átha Cliath Dublin City University Extract 'equivalent width'  $W_{\lambda}$  maps from I & I<sub>0</sub> images.

Here  $\sigma(I)$  is the total absolute photoionization cross section, and the integral is taken over the 46.7 nm (26.5 eV) 5p – 5d resonance profile of Ba+. If ' $\sigma$ ' is known, one can extract the column density NL

 $W_{\lambda} = \Delta \lambda \times \frac{\int I_0(\lambda) d\lambda - \int I(\lambda) d\lambda}{\int I_0(\lambda) d\lambda}$ 

 $W_{\lambda} = \int_{\Omega} \left( 1 - \exp\left( -n_i L \int_{\lambda \lambda} \sigma_{\lambda} d\lambda \right) \right) d\lambda$ 



(A) 100 ns

(B)

Έ

(F)

150 ns

200 ns

300 ns

400 ns

500 ns



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Superpixel





Let's jump forward 20 years to take two recent examples:

- VUV Photoabsorption by Pb & Bi ions [1,2]
- Oxygen K-shell Photoabsorption in CH4 & CO<sub>2</sub> [3]

**1.** *The* 5*d* → 6*p EUV photoabsorption spectra of Pb II and Bi III: evidence of excited states,* H Lu, L Varvarezos, M Bilal Alli, P Nicolosi, J T Costello and P Hayden, J. Phys. B: At. Mol. Opt. Phys. **53** 115001 (2020) - https://doi.org/10.1088/1361-6455/ab7e1c

*Cf. Reply to Comment on* "The 5d  $\rightarrow$  6p photoabsorption spectra of Pb II and Bi III: evidence of excited states" by A. N. Ryabtsev, H Lu, L Varvarezos, M B Alli, P Nicolosi, J T Costello and P Hayden, J. Phys. B: At. Mol. Opt. Phys. **54** Art. No. 068001 (2021) - https://iopscience.iop.org/article/10.1088/1361-6455/abe394

2. The 5d-6p VUV Photoabsorption Spectrum of Bi<sup>+</sup>, H Lu, L Varvarezos, P Hayden, E T Kennedy, J-P Mosnier and J T Costello, Atoms 8, Art. No. 55 (2020) - https://doi.org/10.3390/atoms8030055

3. Soft x-ray photoabsorption spectra of photoionized CH<sub>4</sub> and CO<sub>2</sub> plasmas
 L Varvarezos, H Lu, J T Costello, A Bartnik, P Wachulak, T Fok, L Wegrzynski and H Fiedorowicz, J. Phys. B: At. Mol.
 Opt. Phys. 53 045701 (2020) - https://iopscience.iop.org/article/10.1088/1361-6455/ab5e31





Ollscoil Chathair Bhaile Átha Cliath Dublin City University Motivation - VUV Photoabsorption by Pb & Bi ions

- Spectroscopic observations by the EUVE, Chandra and XMM-Newton satellites emphasise the need for highly accurate atomic data in the extreme VUV (X-VUV) and soft x-ray(SXR).
- Pb II [1] and Bi III [2,3] spectra have attracted attention in the study of stellar evolution and the investigation of the chemical composition of peculiar stars
- Soft x-ray emission measurements in heavy atoms, including Bi and Pb, have been performed in the e.g., the TEXT tokamak [4].
- Generally interested in the short wavelength behaviour of potential metal ions for next generation EUVL resists

[1] Cardelli J A, Federman S R, Lambert D L and Theodosiou C E, 1993 Astrophys. J. **416** L41
[2] Leckrone D S, Johansson S G, Wahlgren G M, Brage T and Profitt C R 1998 Highlights Astron. **11** pp650–652
[3] Wahlgren G M, Brage T, Brandt J C, Fleming J, Johansson S, Leckrone D S, Proffitt C R, Reader J and Sansonetti C J 2001 Astrophys. J. **551** pp520–35
[4] Finkenthal M, Lippmann S, Huang L K, Zwicker A, Moos H W, Goldstein W H and Osterheld A L 1992 Phys. Rev. A 45 pp5846–53





Ollscoil Chathair Bhaile Átha Cliath Dublin City University On short wavelength behaviour of potential metal ions for next generation EUVL (nanoparticle) resists

Related poster: P\_FUN8 - L. Varvarezos - Photoabsorption Spectra of CsI Plasmas in the 50-75 nm Wavelength Range Wednesday – December 1<sup>st</sup>







VUV Photoabsorption by Pb & Bi ions



VUV DLPP Experimental set up. GCA (glass capillary array used for grating protection from plasma debris and as differential pressure barrier).







The 5d  $\rightarrow$  6p EUV photoabsorption spectra of Pb II and Bi III: evidence of excited states, H Lu, L Varvarezos, M Bilal Alli, P Nicolosi, J T Costello and P Hayden, J. Phys. B: At. Mol. Opt. Phys. 53 115001 (2020) - https://doi.org/10.1088/1361-6455/ab7e1c

Comparison of experiment with calculation.





Comparison of experiment with calculation.



**Figure 3.** Calculated gf values for each Bi III configuration involved in our experiment. From top to bottom:  $5d^{10}6s6p^2 \rightarrow 5d^96s6p^3$ ,  $5d^{10}6s^26d \rightarrow 5d^96s^26p6d$ ,  $5d^{10}6s^26p \rightarrow 5d^96s^26p^2$  and  $5d^{10}6s^27s \rightarrow 5d^96s^26p7s$ . A comparison with the experimental spectrum (black line) reveals the contribution of each configuration.

Comparison of experiment with calculation.



Figure 1. The experimentally recorded spectrum of Bi III togethe with the gf values for all the transitions (including 5d<sup>10</sup>6s6p<sup>2</sup> 5d<sup>9</sup>6s6p<sup>3</sup>, 5d<sup>10</sup>6s<sup>2</sup>6d  $\rightarrow$  5d<sup>9</sup>6s<sup>2</sup>6n6d 5d<sup>10</sup>6s<sup>2</sup>6n  $\rightarrow$  5d<sup>9</sup>6s<sup>2</sup>6n<sup>2</sup>



0.12 0.10 Experiment Simulated spectrum 0.08 Bi III 0.06 Ln(l<sub>0</sub>/l) 0.04 0.02 0.00 -0.02 -0.04 50 46 47 49 Wavelength (nm)

**Figure 2.** A comparison between the synthetic spectrum (including  $5d^{10}6s6p^2 \rightarrow 5d^96s6p^3$ ,  $5d^{10}6s^26d \rightarrow 5d^96s^26p6d$ ,  $5d^{10}6s^26p \rightarrow 5d^96s^26p^2$  and  $5d^{10}6s^27s \rightarrow 5d^96s^26p7s$ ) and the experimental spectrum of Bi III.

Table 3. Transitions corresponding to the most prominent features observed in the Bi III spectrum.

Transition	$\lambda_{ m obs}$	$\lambda_{ m calc}$	gf
$5d^{10}6s^26p \rightarrow 5d^96s^26p^2$	(nm)	(nm)	values
${}^{2}P^{o}_{3/2} \rightarrow {}^{2}D_{5/2}$	46.80	46.79	2.47
$^{2}\mathrm{P}_{3/2}^{\mathrm{o}} \rightarrow ^{2}\mathrm{P}_{3/2}$	46.60	46.57	1.28
Transition	$\lambda_{ m obs}$	$\lambda_{ ext{calc}}$	gf
$5d^{10}6s^27s \rightarrow 5d^96s^26p7s$	(nm)	(nm)	values
$^{2}\mathrm{S}_{1/2} \rightarrow  ^{2}\mathrm{P}_{3/2}^{\mathrm{o}}$	45.62	45.58	1.82
${}^{2}S_{1/2} \rightarrow {}^{2}P_{1/2}^{o}$	45.62	45.67	0.85
Transition	$\lambda_{ m obs}$	$\lambda_{ ext{calc}}$	gf
$5d^{10}6s^26d \rightarrow 5d^96s^26p6d$	(nm)	(nm)	values
$^{2}\mathrm{D}_{5/2} \rightarrow ^{2}\mathrm{F}^{\mathrm{o}}_{7/2}$	45.25	45.36	2.06
$^{2}D_{5/2} \rightarrow ^{2}D_{5/2}^{o}$	45.25	45.18	1.20
$^{2}D_{3/2} \rightarrow ^{2}F_{5/2}^{o}$	45.62	45.76	1.55
Transition	$\lambda_{ m obs}$	$\lambda_{ ext{calc}}$	gf
$5d^{10}6s6p^2 \rightarrow 5d^96s6p^3$	(nm)	(nm)	values
${}^{2}\mathrm{D}_{5/2} \rightarrow {}^{2}\mathrm{F}^{\mathrm{o}}_{7/2}$	45.62	45.57	1.15
${}^{4}P_{3/2} \rightarrow {}^{4}P_{5/2}^{o}$	44.96	45.13	1.55
$^{2}\mathrm{P}_{3/2} \rightarrow ^{2}\mathrm{D}_{5/2}^{\mathrm{o}}$	44.60	44.78	1.87
${}^{4}P_{3/2} \rightarrow {}^{2}D_{3/2}^{o'}$	45.62	45.61	0.62
${}^{4}\mathrm{P}_{1/2} \rightarrow {}^{6}\mathrm{D}_{3/2}^{\mathrm{o}}$	44.60	44.63	0.61

Clear evidence of excited states contributing to the overall spectra distribution – difficult to avoid unless they lie significantly higher than plasma temperature – so for more highly charged ions.

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H Lu et al., J. Phys. B: At. Mol. Opt. Phys. 53 115001 (2020)



Comparison of experiment with calculation (the case of Pb II).

Figure 4. The experimentally recorded spectrum of Pb II together with the gf values for all the transitions (including  $5d^{10}6s6p^2 \rightarrow 5d^96s6p^3$ ,  $5d^{10}6s^26d \rightarrow 5d^96s^26p6d$ ,  $5d^{10}6s^26p \rightarrow 5d^96s^26p^2$  and  $5d^{10}6s^27s \rightarrow 5d^96s^26p7s$ ) involved.

**Figure 5.** A comparison between the synthetic spectrum (including  $5d^{10}6s6p^2 \rightarrow 5d^96s6p^3$ ,  $5d^{10}6s^26d \rightarrow 5d^96s^26p6d$ ,  $5d^{10}6s^26p \rightarrow 5d^96s^26p^2$  and  $5d^{10}6s^27s \rightarrow 5d^96s^26p7s$ ) and the experimental spectrum of Pb II.

68

70



Again, clear evidence of excited states contributing to the overall spectra distribution – however simulation is not as successful in this case. Perhaps evidence of missing excited states in the HF basis set expansion or a need to scale the HF-Slater parameters differently to Bi2+ ?



H Lu et al., Atoms 8, Art. No. 55 (2020)



Comparison of experiment with calculation (Bi<sup>+</sup>).

**Figure 3.** A comparison between the synthetic spectra and experimental data. (**a**) Simulated spectrum including photoabsorption from the ground electron configuration of Bi<sup>+</sup> only. (**b**) Simulated spectrum including photoabsorption both from the ground and from low-lying excited states of Bi<sup>+</sup>.

*The 5d-6p VUV Photoabsorption Spectrum of Bi*<sup>+</sup>, H Lu, L Varvarezos, P Hayden, E T Kennedy, J-P Mosnier and J T Costello, Atoms 8, Art. No. 55 (2020) - https://doi.org/10.3390/atoms8030055





Ollscoil Chathair Bhaile Átha Cliath Dublin City University

L Varvarezos et a;, J. Phys. B: At. Mol. Opt. Phys. 53 045701 (2020)



Motivation - Photoabsorption of **Photoionized Small Molecules** 

- Low temperature photoionized plasmas constitute a quite unique state of matter
- Neutral atoms and molecules, radicals, ions and electrons can all be present when radiation in the SXR region is focused onto a dense gas medium
- Unlike the case of plasma formation directly by the optical laser, the SXR photon energy is high enough to ensure the release of core and valence electrons
- Photoionized plasmas are commonly encountered in astrophysics (see for example – B A Remington et al., Rev. Mod. Phys. 78 755–807 (2006)). E.g., plasmas generated in irradiated accretion disks around compact astrophysical objects such as neutron stars, black holes, or white dwarfs
- Quite and interdisciplinary area (J. Phys. D: Appl. Phys. 50 323001 (2017))

Soft x-ray photoabsorption spectra of photoionized  $CH_4$  and  $CO_2$  plasmas, L Varvarezos, H Lu, J T Costello, A Bartnik, P Wachulak, T Fok, L Wegrzynski and H Fiedorowicz, J. Phys. B: At. Mol. Opt. Phys. **53** 045701 (2020) - https://iopscience.iop.org/article/10.1088/1361-6455/ab5e31



- https://iopscience.iop.org/article/10.1088/1361-6455/ab5e31



L Varvarezos et a;, J. Phys. B: At. Mol. Opt. Phys. 53 045701 (2020)

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Photoabsorption of Photoionized Small Molecules

Figure 3. Experimental SXR absorption of the 'not ionized' (a) carbon dioxide and (b) methane plotted together with the normalized oscillator strength adapted from [33]. A gas mixture of xenon/krypton (90/10) was used as the medium to generate the soft x-ray backlighting radiation.

Soft x-ray photoabsorption spectra of photoionized CH<sub>4</sub> and CO<sub>2</sub> plasmas, L Varvarezos, H Lu, J T Costello, A Bartnik, P Wachulak, T Fok, L Wegrzynski and H Fiedorowicz, J. Phys. B: At. Mol. Opt. Phys. **53** 045701 (2020) - https://iopscience.iop.org/article/10.1088/1361-6455/ab5e31



L Varvarezos et a;, J. Phys. B: At. Mol. Opt. Phys. 53 045701 (2020)



Photoabsorption of Photoionized Small Molecules



**Figure 5.** (a) The low-energy part of the Soft x-ray absorption spectrum of  $CO_2$  plotted together with the normalized oscillator strength of CO [33], and calculations performed by means of the Cowan code for C and C<sup>+</sup>. (b) Similar spectrum for CH<sub>4</sub> plotted together with the calculated absorption spectrum of CH<sub>3</sub> [41], and calculations performed by means of the Cowan code [46, 47] for C and C<sup>+</sup>.



L Varvarezos et a;, J. Phys. B: At. Mol. Opt. Phys. 53 045701 (2020)



#### Photoabsorption of Photoionized Small Molecules

**Table 2.** Tabulated transitions for the  $CO_2$  and  $CH_4$  plasmas, including transitions in  $CH_3^+$ , CO, C, C<sup>+</sup> species.

Photon energy (eV)	Molecular/ atomic species	Assignment	
(eV) 281.8 282.2 284.2 284.2 284.3 284.5 285.9 286.2 287.1 287.2 287.4 287.7 287.8 287.9 287.9 288.2	species CH <sub>3</sub> <sup>+</sup> CH <sub>3</sub> <sup>+</sup> CI CI CI CI CI CII CII CII CI	Assignment $(1\alpha'_{1})(C1s) \longrightarrow (1\alpha''_{2})(0-0)$ $(1\alpha'_{1})(C1s) \longrightarrow (1\alpha''_{2})(1-0)$ $1s^{2}2s^{2}2p^{2}(^{1}S)^{3}P \longrightarrow 1s^{1}2s^{2}2p^{3}(^{4}S)^{3}S$ $1s^{2}2s^{2}2p^{2}(^{1}S)^{3}P \longrightarrow 1s^{1}2s^{2}2p^{3}(^{2}D)^{3}D$ $1s^{2}2s^{2}2p^{2}(^{1}S)^{1}D \longrightarrow 1s^{1}2s^{2}2p^{3}(^{2}D)^{1}D$ $1s^{2}2s^{2}2p^{2}(^{1}S)^{1}D \longrightarrow 1s^{1}2s^{2}2p^{3}(^{2}P)^{1}P$ $1s^{2}2s^{2}2p^{2}(^{1}S)^{1}D \longrightarrow 1s^{1}2s^{2}2p^{3}(^{2}P)^{1}P$ $1s^{2}2s^{1}2p^{2}(^{4}S)^{4}S \longrightarrow 1s^{1}2s^{1}2p^{3}(^{3}P)^{4}P$ $1s^{2}2s^{1}2p^{2}(^{2}D)^{2}D \longrightarrow 1s^{1}2s^{1}2p^{3}(^{3}P)^{2}P$ $1s^{2}2s^{1}2p^{2}(^{2}D)^{2}D \longrightarrow 1s^{1}2s^{1}2p^{3}(^{3}P)^{2}P$ $1s^{2}2s^{2}2p^{1}(^{1}S)^{2}P \longrightarrow 1s^{1}2s^{2}2p^{2}(^{1}D)^{2}D$ $1s^{2}2s^{2}2p^{1}(^{1}S)^{2}P \longrightarrow 1s^{1}2s^{2}2p^{2}(^{3}P)^{2}P$ $1s^{2}2s^{1}2p^{2}(^{2}D)^{4}D \longrightarrow 1s^{1}2s^{1}2p^{3}(^{3}P)^{4}P$	For photoionized CO2, oxygen absorption of Oxygen K-shell photoabsorption spectra of photoionized CO2 plasmas L Varvarezos, Hu Lu, J T Costello, A Bartnil Wachulak, T Fok, Ł Węgrzyński and H Fied J. Phys. B: At. Mol. Opt. Phys. <b>53</b> 105701 (2 https://doi.org/10.1088/1361-6455/ab78ab
287.4	CO	$2\sigma(C1s) \longrightarrow 1\pi_u, \pi \text{ resonance}$	

Soft x-ray photoabsorption spectra of photoionized  $CH_4$  and  $CO_2$  plasmas, L Varvarezos, H Lu, J T Costello, A Bartnik, P Wachulak, T Fok, L Wegrzynski and H Fiedorowicz, J. Phys. B: At. Mol. Opt. Phys. **53** 045701 (2020) - https://iopscience.iop.org/article/10.1088/1361-6455/ab5e31

#### **DLPP - Some final points**

- Opens up both valence and inner shell excitation of atoms and molecules inn neutral to moderately ionized states
- Offers a degree of control over the ion stage under study (dx, dt/ laser-I,  $\lambda$ , etc)
- Opens up the study of photoabsorption along isonuclear and isoelectronic sequences
- Has the possibility to track and extract 'column densities' in space and time
- Can complement optical diagnostics sense cool environments
- Can complement/augment work on merged ion beams with synchrotron experiments (e.g., [1]) or multiple photoionization with FELs [e.g., 2])
- 1. The photon-ion merged-beams experiment PIPE at PETRA III The first five years, S Schippers et al., X-Ray Spectrometry **49** pp11–20 (2020)
- Near-Threshold Two-Photon Double Ionization of Kr in the Vacuum Ultraviolet, L Varvarezos et al, Phys. Rev A 103 Art. No. 022832 (2021)

L Varvarezos et a;, J. Phys. B: At. Mol. Opt. Phys. 53 045701 (2020)

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### Perspectives for X-VUV laser/laser plasma based research



Ollscoil Chathair Bhaile Átha Cliath Dublin City University

L Varvarezos et a;, J. Phys. B: At. Mol. Opt. Phys. 53 045701 (2020)

- Will mainly remain a 'niche' research area but works where offline is possible
- Benefits of access to resonance lines and continua of all atomic species remain
- New frontiers e.g., attosecond transient absorption spectroscopy (ATAS) [1]
- Non linear photoionization processes involving inner shells. E.g., two X-VUV photon absorption leading to femtosecond photo and Auger electron emission [2,3].
- Pump-probe with fs jitter (using injection seeded X-VUV FELs) [4,5]
- 1. UF dissociation of vinyl bromide: An attosecond transient absorption spectroscopy and non-adiabatic molecular dynamics study, R Flott et al., Struct. Dyn. **8**, 034104 (2021) / J. Phys. B: At. Mol. Opt. Phys. 49 (2016) 062001
- 2. Two-photon inner-shell ionization in the extreme-ultraviolet, V Richardson et al., Phys. Rev. Lett 105, Art. No. 013001 (2010)
- 3. Clocking Auger electrons, D C Haynes et al., Nature Physics 17 pp.512–518 (2021)
- 4. Coherent control with a short-wavelength free-electron laser, K. C. Prince et al., Nat. Phot. 10, 176 180 (2016)
- 5. Tracking attosecond electronic coherences using phase-manipulated extreme ultraviolet pulses, A. Wituschek et al., Nature Communications volume 11, Article No. 883 (2020)



#### Perspectives for X-VUV laser/laser plasma based research



Free-electron laser pulse

L Varvarezos et a;, J. Phys. B: At. Mol. Opt. Phys. 53 045701 (2020)



Neon

target











FIG. 2. Photoelectron yields for a range of time delays in the vicinity of the half revival period.



Nature Physics 17 pp.512–518 (2021)

Fig. 1 | Mid-infrared streaking. 17 µm mid-infrared (MIR) streaking

(OPA) and difference frequency generation (DFG), and coupled into a

with 7 fs, 1,130 eV free-electron laser pulses in a neon gas target. The

large-acceptance time-of-flight (TOF) spectrometer.

resultant streaked photo and Auger electron emission is measured using a

chamber through a potassium bromide (KBr) window. The mid-infrared pulses are focused with a parabola of focal length 100 mm and overlapped

laser pulses are generated by downconversion of a near-infrared

#### Perspectives for X-VUV laser/laser plasma based research



## TA – LIBS in the UV—Visible-NIR?

# Current (TA)-LIBS Focus and Future Developments



Focus: Wind Energy – Wind Turbine Blades

NATURE | VOL 412, 2001, pp42-43





Spectrochimica Acta Part







#### EUVL Resist Materials (Synthesis & Dynamics) -

# Current (TA)-LIBS Focus and Future Developments



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# Related Poster - P\_IND6 - S. Cummins - UV-Visible and Vacuum-UV LIBS on Wind Turbine Blade Samples Wednesday – December 1<sup>st</sup>





# **Current LIBS Focus and Future Developments**





#### Theme: *LIBS for Future Pharmaceutical PATs* Focus: ML- Assisted Chemometrics (P Hayden (now UCD) + IBM)

M B Alli, PhD Thesis, DCU (2021)

Table 5.4: % Accuracy ,comprehensive study, 160 spectra tested.

Wavelength Range					
Technique	65 nm	105 nm	Average		
SOM 2x2	52.50	0.00	26.25		
SOM 3x3	81.25	10.63	45.94		
SOM 4x4	73.13	32.50	52.81		
SOM 5x5	81.25	13.75	62.50		
PCA	66.25	41.88	54.06		
COMP	69.38	0.63	35.01		
CNN	95.00	91.88	93.44		



# **Current LIBS Focus and Future Developments**



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#### Theme: Light Elements in Steel

Focus: Line focus VUV LIBS (and Time Resolved LIBS)

Related Poster - P\_CQ8 – S S Zehra, L Varvarezos et al. - Limit of Detection for Carbon in Steel using Time Integrated Space Resolved (TISR) VUV LIBS: Line versus Point Plasma Plumes.

Wednesday – December 1<sup>st</sup>



## **Current Students/PDs**

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SEAI - Sustainable Energy Authority Of Ireland- Research, Development & Demonstration(RD&D) programme

